Calibration for GEMS Data, Concepts and Issues

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4th GEMS Science Meeting (2013.10.14~10.16.)

Data quality of the GEMS

Radiometric calibration and issues

INR and Issues



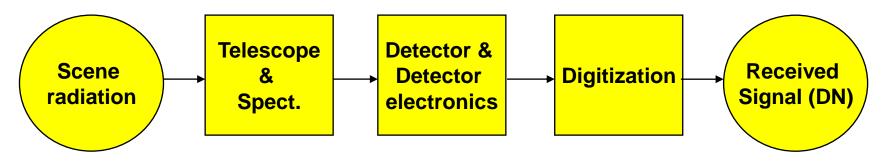
I. Data quality

Determined by three important processes

– Spectral Calibration

- Assignment of wavelength value to each CCD pixel (the 1st order spectral calibration will be prepared by instrument manufacturer)
- For a better accuracy, additional steps could be taken
- Radiometric Calibration
 - Convert the downlinked digital number to physical parameter (here, either radiance or irradiance)
 - Need comprehensive ground test and onboard operations for the confirmation or re-characterization
- Geometric Calibration
 - Assignment of geo-location information to each CCD pixel and corresponding observation angles (solar/satellite zenith angle,)
 - Dependent not only on instrument characterizations, but also on the spacecraft and ground processing
 - Need early test for the landmark performance with the reduced spatial resolution data

Signal chain equation



 Then, the digital number received at the ground is determined by;

$$DN = C \times \left\{ G\left[\int_{\nu_1}^{\nu_2} I(\nu) T(\nu) \frac{D^2 A}{f l^2} R(\nu) d\nu + \varepsilon_0 \right] + \varepsilon_e \right\}$$
(1)

here,

DN: digital number, C: conversion constant, G: amplification gain, I(v): input radiance, T(v): optical transmittance D: Aperture diameter, A: detector area, fl: focal length R(v): detector spectral responsivity v: wavenumber, ε_o : opitcal error, ε_e : electronic error

◆ Calibration equation

– Chain equation can be inverted to give a band-averaged input radiance (\overline{I}) value such as;

$$\bar{I} = \frac{1}{G'C} \frac{1}{\int_{v_1}^{v_2} T(v) R(v) dv} \{DN - C\varepsilon_e - CG\varepsilon_o\}$$

$$= \frac{k}{\tau} DN' \qquad (2)$$
where
$$k = \frac{1}{G'C},$$

$$\tau = \int_{v_1}^{v_2} T(v) R(v) dv,$$

$$DN' = DN - C\varepsilon_e - CG\varepsilon_e$$

– Here, k is ground determined and time invariant, while R(v) weighted T(v) (or called throughput) and error terms ($\varepsilon_e \& \varepsilon_o$) are all time dependent

◆ Calibration equation

- As the throughput degrades/varies with time, the normalized radiance (w.r.t to the simultaneous solar irradiance value) has long been used
- Similar to the radiance calibration equation, irradiance calibration equation can be denoted as;

$$F = \frac{k_s}{\tau} D N_s' \frac{1}{g \widetilde{T}}$$

- Here, g is the goniometry (angular response to the solar radiation, and \tilde{T} is the transmissivity of the transmissive diffuser used for irradiance observation
- Then, the normalized radiance becomes

$$\frac{I}{F} = \frac{k}{k_s} \frac{DN'}{DN_{s'}} g\tilde{T}$$
(3)

◆ Key issues with the radiometric calibration

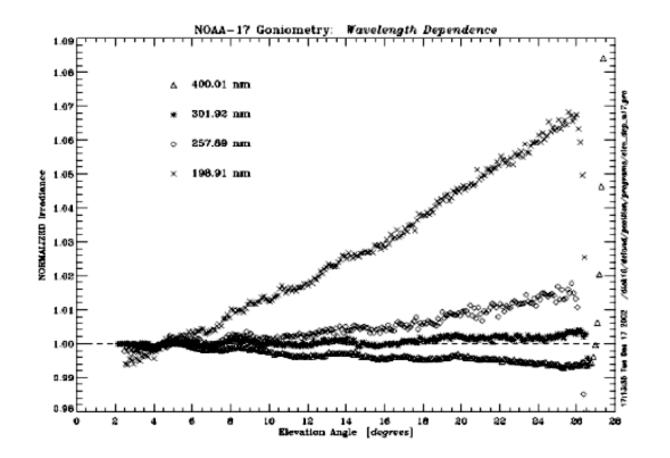
$$\frac{I}{F} = \frac{k}{k_s} \frac{DN'}{DN_{s'}} g\tilde{T}$$
(3)

- Acquiring accurate ground measured calibration parameters (calibration constants, goniometry, non-linearity, CCD alignment characteristics, baseline solar irradiance, etc.)
- Characterization of noise counts for both optics and electronics from the on-orbit observation data
- Any degradation of the goniometry parameter with time?
- How can we best utilize the frequent observation of solar radiation (proposed for twice a day)
- Do we need additional normalization to control PRNU within a certain threshold value?

◆ Solar radiation for absolute calibration

- Use transmissive diffuser to measure solar irradiance twice a day
 - Are the measured solar irradiance going to be used for the derivation of the normalized radiance?
 - What time are we going to measure the solar irradiance?
 - Does GEMS have enough FOR to have 0.53 degree solid angle sun without satellite tilting?
 - How are these frequent solar irradiance data going to be used?
- Then, how often the reference diffuser should be used?
- Are we going to use the reference and working diffuser with the similar concept applied for the OMPS nadir spectrometer?
- How accurately can we measure the goniometry parameter at the ground testing? Can IOT tests provide better information?

◆ Goniometry effect



L. Flynn et al, (2011)

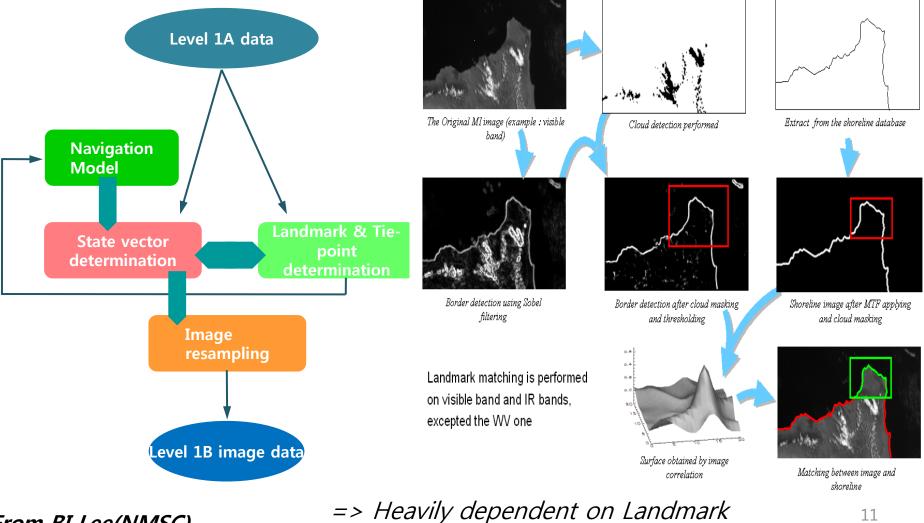
How accurately can we correct the background noise

- Are we going to have a tool to correct the stray light effect in the measured radiance?
 - Will depend on the wavelength and scene type. How are we going to implement laboratory obtained information to the real data?
- Striping caused by the PRNU and other reasons are observed quite often. What kind of protective measures do we have/expect to have or do we anticipate?
- How exactly are we going to measure the LED signal and use the signal for the calibration (linearity trending)?
- Smearing effect would depend on the observation environment. What kind of correction measures are we going to have?

III. INR and Issues

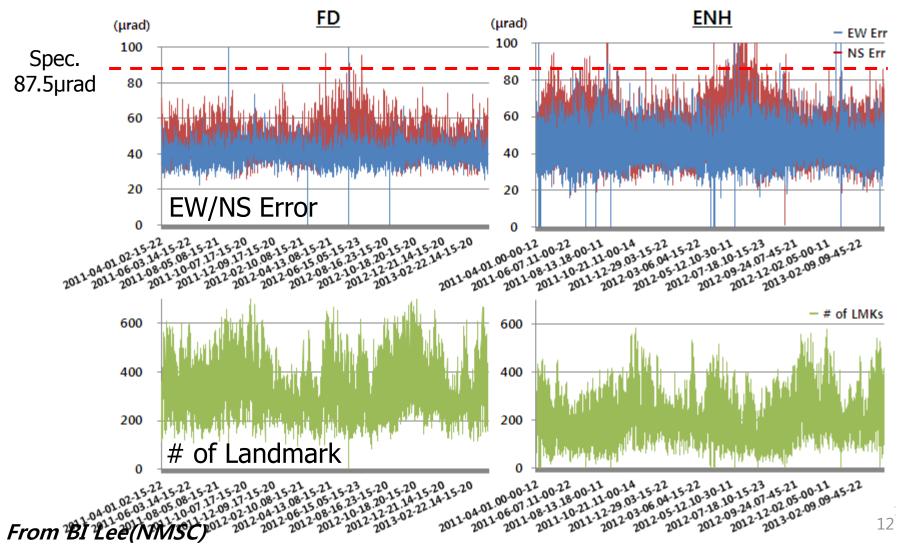
COMS INR

From BI Lee(NMSC)



III. INR and Issues





III. INR and Issues

◆ GEMS Issues with INR

- Slow scan rate with relatively small FOR
 - Reduces available number of good landmarks within the FOR
 - Require more uniformly distributed good landmarks and better ranging information
- Spatial resolution
 - Increased possibility of cloud contaminated pixel which could decrease available number of landmark.
 - Lower spatial resolution means a larger landmark size (to have same number of coastal pixel number) resulting decreased number of landmark
- Limited number of observation
 - The first scene without recent state vector information (such as the first scene of a day, or, after spacecraft maneuvering such as station keeping and wheel-off loading)
 - Limited portion of spectral data will be used for Landmark extraction, which may introduce spectral mis-alignment

IV. Summary

- Many aspects are new for the geostationary platform
- ◆ Interesting and exciting activities are ahead
- ◆ Need closer cooperation
 - Image quality is the most important for the better utilization of GEMS data
 - It is determined by at least three important calibration activities, Geometric, Spectral, and Radiometric.
 - All of these calibration require a closer collaboration of all stakeholder including user, sensor/spacecraft engineer, ground processing and scientific algorithm developer

Thank you