### **Activities at Max Planck Institute for Chemistry for Sentinel 4**

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- Specific challenges of verification activities for S4
- Products and involved institutions
- Generation of test data sets (TDS) (In cooperation with S4 team)



# **Sentinel 4/UVN Instrument**

- Geostationary satellite monitoring Europe
- MTG-Sounder (MTG-S) satellite
- 2D hyperspectral detector
- Resolution: 0.5 nm (UV-VIS), 0.12 nm (NIR)
- Spectral range:



• Launch planed in 2021







Stark, ESA

# S4 scanning and data acquisition

- **GCA** = Geographic Coverage Area
- $=$  Reference Area RA
- = Instantaneous Field Of View **IFOV**

#### **OZA = observation zenith angle**



- SSD  $\approx$ 8 km in E/W and in N/S at ref point (45N lat, lon of SSP)
- Angular sampling grid ~evenly spaced in Field Angle and Scan Angle ٠
- SSD varies spatially across GCA with distance sat-target and VZA at target ٠

Graphic & text adapted from Ben Veihelmann, ESA

#### **Specific challenges of the S4 mission**

Because S4 is ESA's first geostationary atmospheric chemistry mission, no measurement data with the same properties can be used.

 $\triangleright$  Verification activities focus on synthetic data!

Diurnal and seasonal variations of atmospheric properties have to be covered:

- trace gases
- Aerosols
- clouds

Realistic surface reflectance data (BRF) have to be considered

#### **Specific challenges of the S4 mission**

ESA's first geostationary atmospheric composition mission

- $\triangleright$  the simulation and interpretation of the atmospheric radiative transfer is different compared to previous missions in many respects:
- The same ground scenes will be observed under varying solar illumination conditions (SZA and relative azimuth angle) during the course of the day (and also for different seasons).
- BRF-Effects (for trace gas, but also for cloud and aerosol products) have to be considered
- Limited spatial coverage: new , reference regions' and reference techniques have to be found for trace gas retrievals (e.g. to estimate stratospheric NO2)

#### **Level-2 Processor Development Consortium lead by DLR**





• Simulation of Synthetic spectra (Test Data Set) at MPIC

## **Generation of synthetic spectra**

### **Approach 1: Set of selected locations**

- full diurnal cycle of three consecutive days both for June and December 2003
- the focus of this approach is to represent *temporal variability* over the day
- 3x3 pixel area around selected S4 pixel



Selected pixel 3x3 area around S4 pixel grid

### **Approach 2: whole S4 area**

- two days in June and December 2003.
- the focus of this approach is to cover the *spatial variability*.
- only for every 100th S4 pixel a full spectrum will be calculated (by selecting every 10th pixel in each direction)

### **Selected Pixel Approach 1**



#### **Approach 2**



Example for the sampling of the full S4 scan for approach 2.

#### **Input into simulations**

- 1) **NO<sup>2</sup> , HCHO and O<sup>3</sup>** : Diurnal variations of trace gas concentration profiles are taken from model simulations:
- 0 to 3.5km: Lotos-EUROS, high spatial resolution (MACC domain (15°W-35°E, 35°N-70°N) with about 7 km horizontal resolution)
- above 3.5km and outside MACC domain: TM5, lower spatial resolution (1°x1° , covering the area from 80°W to 85°E and 14 $\degree$ N to 86 $\degree$ N) (model data from Isotrop project, Henk Eskes pers. comm.)

2) **SO<sup>2</sup>** : prescribed scenarios for Boundary layer pollution and volcanic emissions are used



Variation of the vertical grid of the LOTOS-EUROS model (left) and the TM5 model (right) as function of the latitude for a longitude around 2.5°. The lines at the lowest altitude indicate the surface elevation (data for 01 June 2003, 9:00 UTC).



 $\mathsf{NO}_2$  surface mixing ratio over Europe from LOTOS-EUROS (left) and TM5 (right). (data for 01 June 2003, 9:00 UTC).



 $NO<sub>2</sub>$  mixing ratios at different altitudes. Left: Surface mixing ratio from combined LOTOS-EUROS and TM5 model data. Right mixing ratio at ~24.5km altitude from TM5. (data for 01 June 2003, 9:00 UTC). The sharp gradient in the west is caused by the terminator.

#### **SO<sup>2</sup> :**

#### **A) pollution scenarios:**

Two SO<sub>2</sub> VCDs will be used, representing low  $(1.10^{16} \text{ molecule})$  and moderate  $(5.10^{16}$  molec/cm<sup>2</sup>) pollution levels. In recent years SO<sub>2</sub> over Europe has decreased significantly



SO<sub>2</sub> VCDs over the middle east derived from OMI observations (Christoph Hörmann, MPIC). The VCDs were determined using a geometric AMF. Thus the true  $SO<sub>2</sub>$  VCDs are probably by a factor 2 to 4 larger.

#### **B) Volcanic eruptions and degassing**

Plumes will be simulated for two scenarios:

- emissions from continuously degassing volcanoes:
	- moderate  $SO<sub>2</sub> VCD (1.10<sup>17</sup> molec/cm<sup>2</sup>)$
	- in a layer between 2 and 3 km.
- volcanic eruptions:
	- moderate  $SO<sub>2</sub> VCD(1.10<sup>17</sup> molec/cm<sup>2</sup>)$
	- high  $SO<sub>2</sub> VCD (1.10<sup>19</sup> molec/cm<sup>2</sup>)$
	- in a layer between 8 and 9 km.

(These settings match those used within S5P verification)



C. Hörmann, MPIC



C. Hörmann, MPIC

#### **3) Aerosols:**

- AOD is taken from average MODIS observations (see below); in addition to these values also the half and double values are used.
- Angstrom parameter is set to 1
- the profile height is taken from the model data (boundary layer height)
- for desert dust aerosols an elevated aerosol layer between 3 and 4 km altitude with an AOD of 1 is assumed





Average AOD at 550 nm derived from MODIS observations (2003-2016).

© Giovanni online data system, developed and maintained by the NASA GES DISC)

### **4) Clouds:**

- typical cloud scenarios (clear, partly cloudy, totally overcast) will be assumed
- cloud heights: 1-2km, 3-4 km, and 7-8km
- cloud OD: 5 and 20
- the independent pixel approximation (IPA) will be used

#### **5) Temperature and pressure profiles**

• taken from the TM5 model

#### **6) Surface BRF data**

- Surface BRF data from the GRASP algorithm
- on land: BRF kernels are based on the Ross-Li parameterisation (3 kernels obtained from POLDER observations)
- over ocean: Cox-Munk parameterisation

#### **7) Surface elevation**

 $\triangleright$  To be consistent with the model data surface elevation is taken from the model data LOTOS-EUROS & TM5

#### **BRF parameters for a selected site (Cairo)**



MODIS image of Cairo; the red polygon indicates the S4 ground pixel.



Geometries of observation (red) and illumination (blue) over the Cairo site in summer.



BRF kernels from the Ross-Li parameterisation for Cairo in Summer as function of the wavelength.



Surface BRF simulated over the Cairo site with the Ross-Li model in the 5 selected wavelengths.

#### **Generation of test data sets (TDS)**

#### • **High resolution spectra**:

They will be provided to ESA as input to S4/UVN instrument simulations and for the calculation of synthetic L1b data. High resolution spectra will also be provided to project partners for specific tasks, e.g. to investigate the effect of heterogenous scenes on the L1b spectra.

• **S4 spectra**: spectral resolution (L1b) of the S4 instrument. They are derived from high resolution spectra by convolution and

resampling according to the spectral properties of the S4 instrument.

Noise will be calculated based on the simulated radiance and the instrumental properties

#### **Simulation details:**

- Simulations performed with SCIATRAN RTM
- Viewing geometry (SZA, VZA, rAA) according to time and location of selected ground pixel
- BRF according to time and location of selected ground pixel
- profiles of trace gases, aerosols and clouds are taken from model simulations or prescribed scenarios

Trace gas cross sections:

UV/VIS : NO2, HCHO, O3, H2O, O4, SO2 NIR: O2, H2O

- Rotational Raman is considered\*
- polarisation is considered\*

Sampling and resolution UV/VIS:

- high Resolution spectra: sampling: 0.01 nm, resolution: 0.04 nm
- low resolution spectra: sampling: ~0.16 nm, resolution: 0.5 nm

Sampling and resolution NIR :

- high resolution spectra: sampling 0.001 nm
- low resolution spectra: sampling: ~0.04 nm, resolution: 0.12 nm \*first data set without polarisation and Raman scattering

### **Examples of UV/VIS high and low resolution spectra**



Milano (45.47°N 9.18°E, ~12:00 UTC, Raman scattering & polarization included

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### **Examples of NIR high and low resolution spectra**



Milano (45.47°N 9.18°E, ~12:00 UTC, Raman scattering & polarization included

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### **Summary**

 Verification activities are crucial for testing the operational algorithms

- For S4 (as the first geostationary chemistry mission) the verification activities are based on synthetic data
- MPIC is responsible for the definition and generation appropriate test data sets
- All relevant atmospheric scenarios and viewing geometries have to be covered
- Bidirectional reflectance properties of the surface have to be considered
- Spectra are simulated with high spectral resolution and considering polarisation and Raman scattering