

# **GEMS**

## **Polarization Correction**

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# Background

- ❑ Sunlight is polarized when reflected from the earth-atmosphere system.
- ❑ Radiometric response of an instrument depends on the polarization of the incoming light (Schutgens and Stammes, 2003).
- ❑ To reduce the instrument polarization sensitivity, two methods are used.
  - Depolarization method
    - destroys the polarization information by scrambling
    - used by TROPOMI, OMI, TOMS, SBUV
  - Polarization characterization method
    - characterizes instrument polarization sensitivity and atmospheric polarization
    - used by GOME, GOME-2, SCIAMACHY

# Background

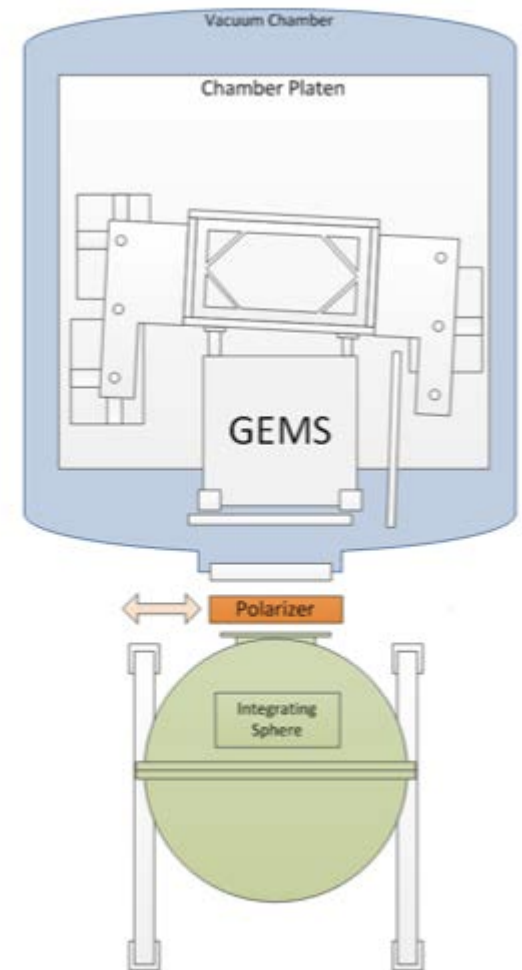
- ❑ **Some instruments measure the state of polarization primarily for the purpose of improving their radiometric calibration.**
  - GOME (Burrows et al., 1999)
  - SCIAMACHY (Bovensmann et al., 1999)
  - GOME-2 (Callies et al., 2000)
  
- ❑ **GEMS does not have a sensor that observes polarization state.**
  - GEMS will use a polarization correction algorithm based on RTM simulation results.
  - Enables a more accurate retrieval of atmospheric properties and constituents.

# GEMS Polarization Ground Test

- ❑ A wire-grid polarizer is placed in the illumination path.
- ❑ The polarizer rotates from 0° to 725°.  
(5° interval)
- ❑ LPS (Linear Polarization Sensitivity) and PA (Polarization Axis) are derived.

$$LPS = \frac{I_{max} - I_{min}}{I_{max} + I_{min}}$$

- ❑ The model results (LPS ratio, from BATC) are applied to LPS and PA at the center of N/S and E/W scan mirror positions.



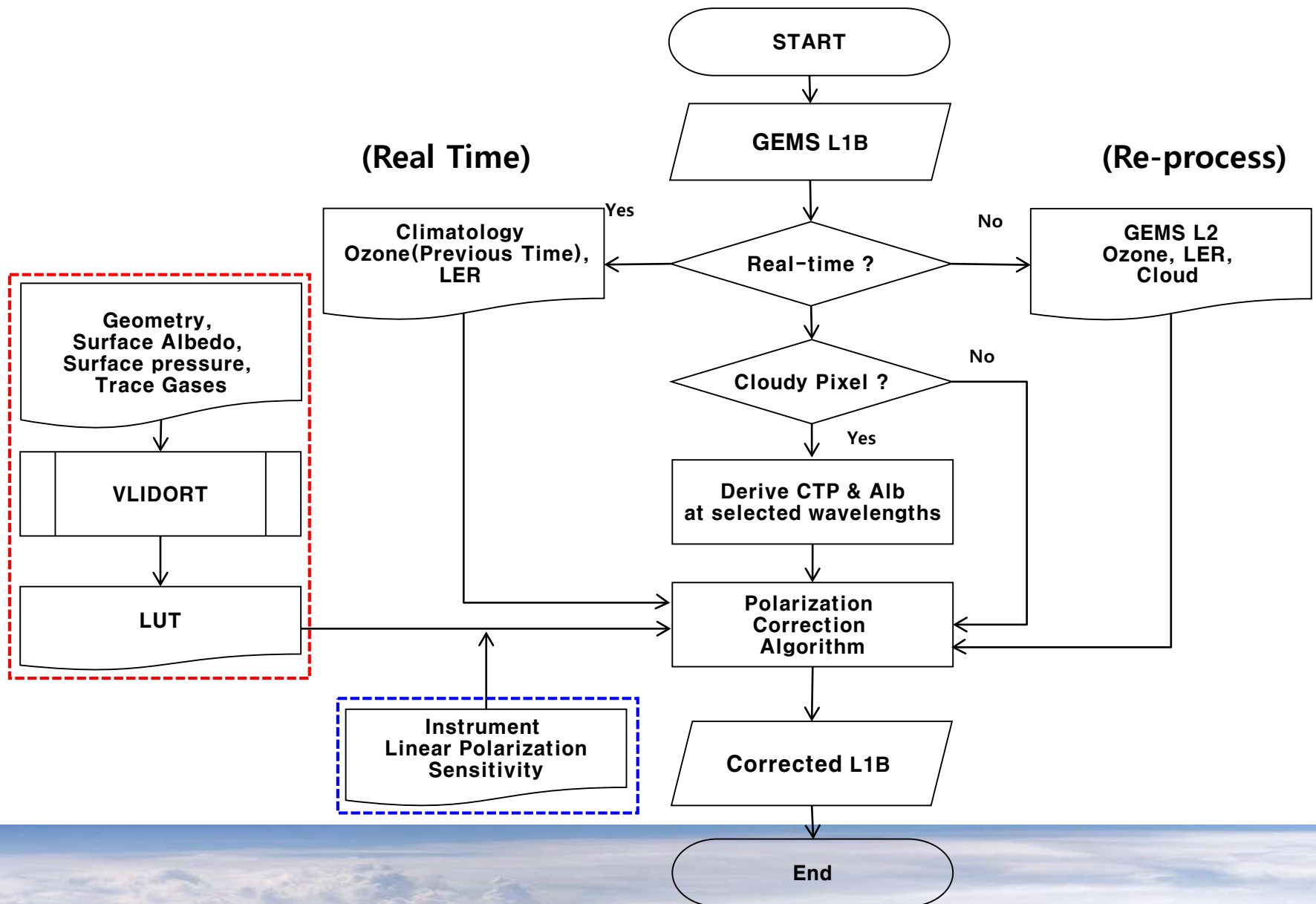
Calibration Test Station (CATS)

# GEMS Linear Polarization Sensitivity

## User Requirements

- ❖ Less than 2 %.
  - ❖ No inflection point within 20 nm wavelength range.
- 
- ❖ Considerable changes of LPS and PA were reported.
  - ❖ Requirements are not satisfied in some regions.

# Flow Chart



# Polarization Correction Algorithm

## □ Polarization Correction Algorithm (Sun and Xiong, 2007)

$$I' = hI\{1 + \underbrace{f a \cos[2(\phi - \chi)]}_{\text{Polarization Correction Term}}\}$$

Diagram annotations: A blue arrow labeled "instrument" points to the variable  $f$ , which is circled in blue. A red arrow labeled "Atmosphere" points to the variable  $a$ , which is circled in red. A red bracket underneath the entire term  $f a \cos[2(\phi - \chi)]$  is labeled "Polarization Correction Term".

$I'$ : GEMS L1B (Mesarused)

$h$ : Transmittance (Radiometric calibration coefficient; assume to 1)

$I$ : True Intensity (Corrected L1B)

$a$ : Degree of (linear) polarization

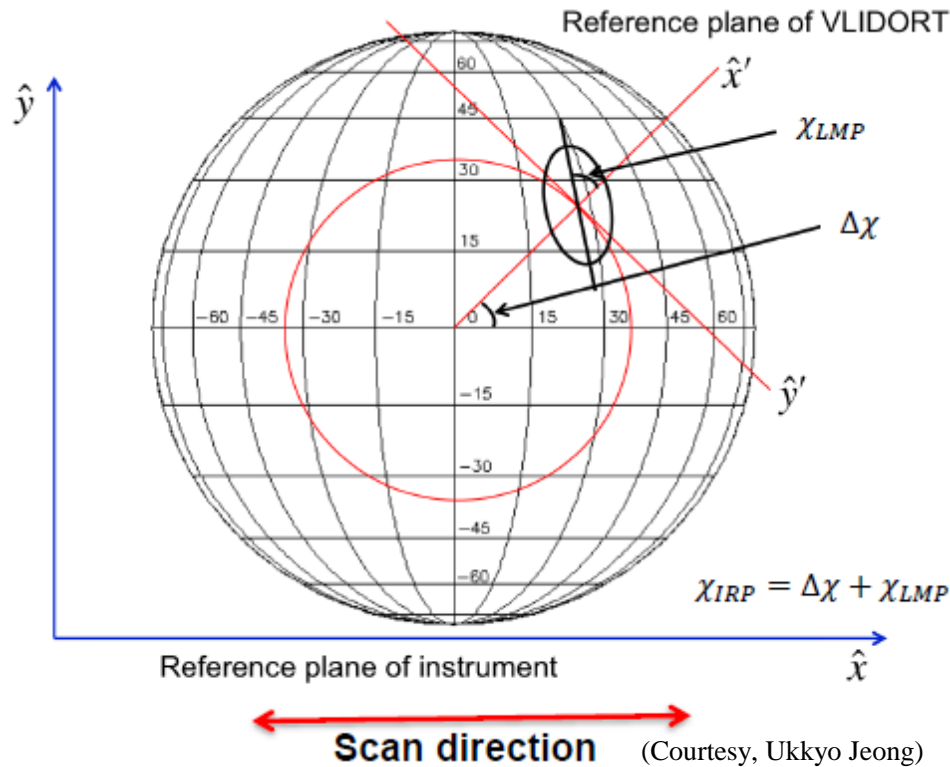
$\chi$ : Polarization Axis

$\phi$ : Angle of polarization w.r.t. instrument reference plane

$f$ : Linear Polarization Sensitivity (GEMS Polarization Factor)



# Polarization Angle



$$\chi_{LMP} = \frac{1}{2} \arctan\left(\frac{U}{Q}\right)$$

$$\Delta\chi = \tan^{-1} \left[ \frac{\sin\theta}{\cos\theta \sin(\Delta\phi)} \right]$$

$$\chi_{IRP} = \chi_{LMP} + \Delta\chi$$

$\chi_{LMP}$ : Angle of polarization w.r.t. Local Meridian Plane (LMP); calculated by VLIDORT

$\chi_{IRP}$ : Angle of polarization w.r.t. Instrument Reference Plane (IRP)

$\Delta\chi$ : Difference of polarization angles for IRP and LMP

$\theta$ : Latitude of ground location

$\Delta\phi$ : Difference of Longitude between Satellite and ground location

# Look-Up Table

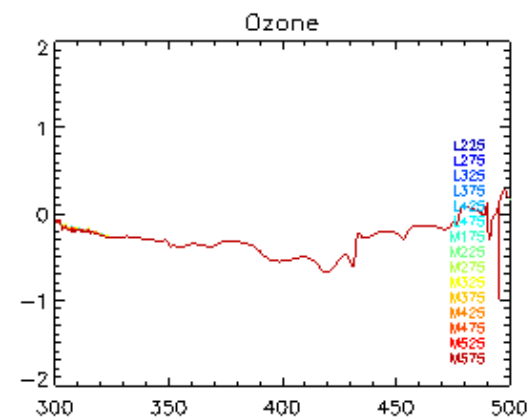
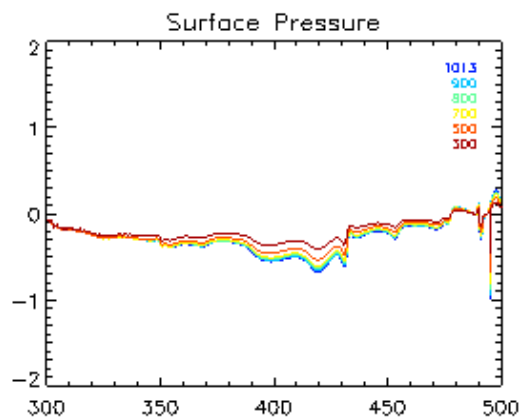
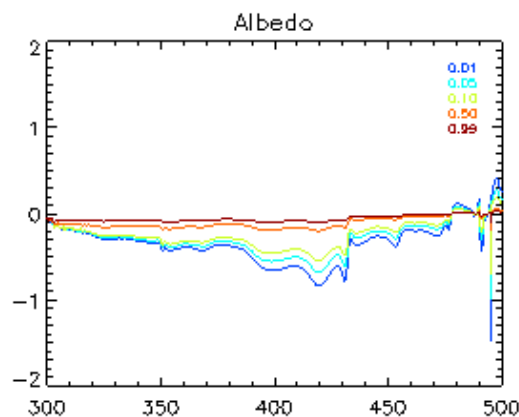
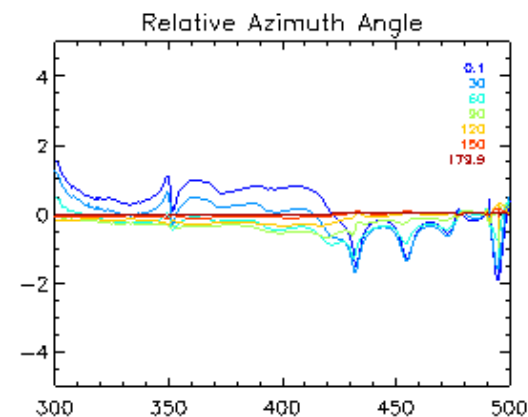
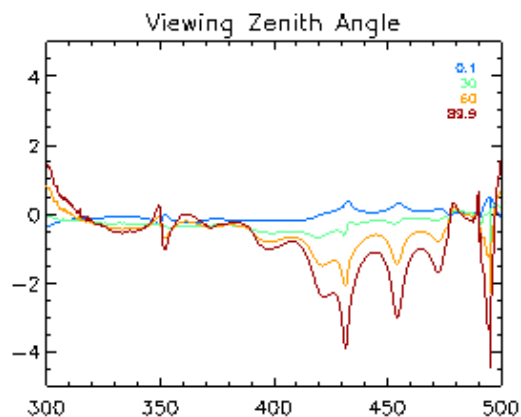
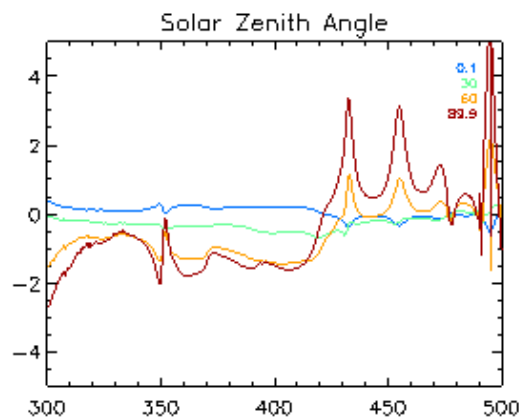
❖ Atmospheric Stokes Parameters(I, Q, U) are calculated using VLIDORT as a function of SZA, VZA, RAA, Albedo, Surface pressure, and ozone.

- US76 standard atmosphere with O<sub>3</sub>, NO<sub>2</sub>, SO<sub>2</sub>, HCHO, O<sub>2</sub>-O<sub>2</sub>
- Ozone Profiles are based on TOMS V8 climatology
- Rayleigh scattering

| Parameter   | Nodes   |
|---|---|
| Spectral Resolution [nm]                          | $\Delta 0.2$ ( 300 ~ 500 )  |
| SZA [degree] (10)                                 | 0.1, 10, 20, 30, 40, 50, 60, 70, 80, 89.9   |
| VZA [degree] (10)                                 | 0.1, 10, 20, 30, 40, 50, 60, 70, 80, 89.9   |
| RAA [degree] (11)                                 | 0.1, 5, 30, 45, 60, 90, 120, 135, 150, 175, 179.9   |
| ALBEDO (5)  | 0.01, 0.05, 0.10, 0.50, 0.99  |
| Surface pressure [hPa] (12)                       | 1013, 900, 800, 700, 500, 300, 200  |
| Ozone profiles [DU] (21)<br>0 ~ 30(L), 30 ~ 60(M) | M175, M225, M275, M325, M375,<br>M425, M475, M525, M575<br>L225, L275, L325, L375, L425, L475 |

# Polarization Error Sensitivity

$$\text{Relative Difference [\%]} = \frac{I' - I}{I} * 100$$



## Base Condition

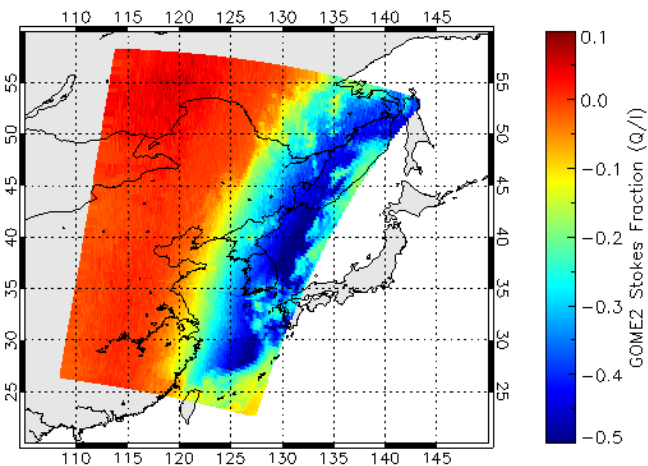
- Ozone : M325
- Alb : 0.05
- Surface Pressure : 1013 hPa
- SZA : 30
- VZA : 30
- RAA : 90

**Geometry >> Alb > Surf\_pres > O3**

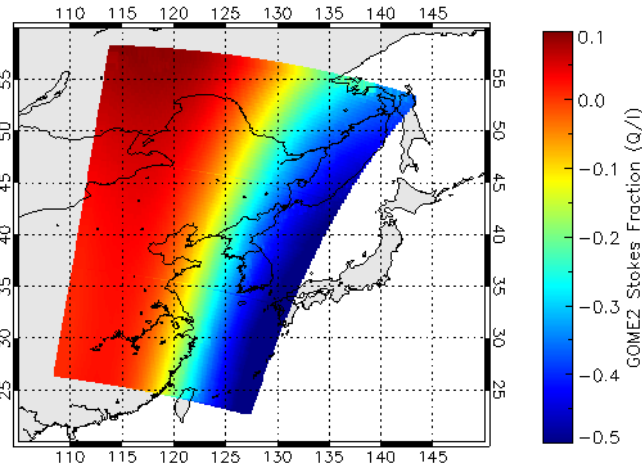
※ Without Polarization Correction,  
Radiance errors are up to 2 %

# Verify RTM simulation (w/ GOME -2 PMD)

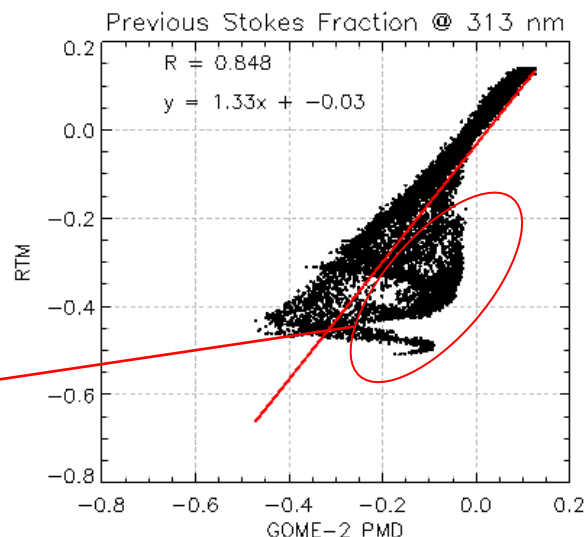
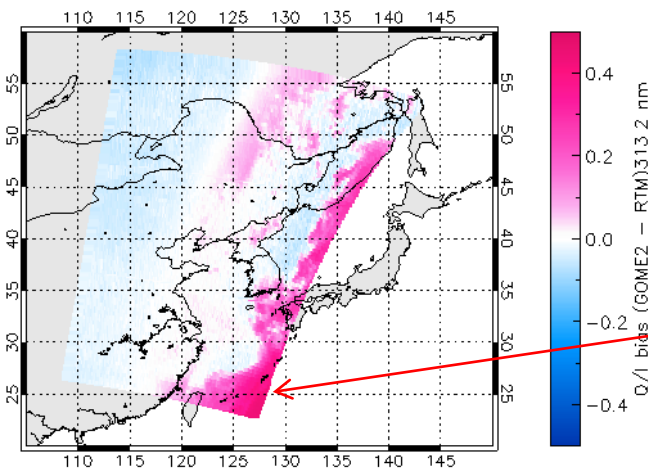
GOME-2 PMD @ 313.2 nm



RTM results @ 313.2 nm

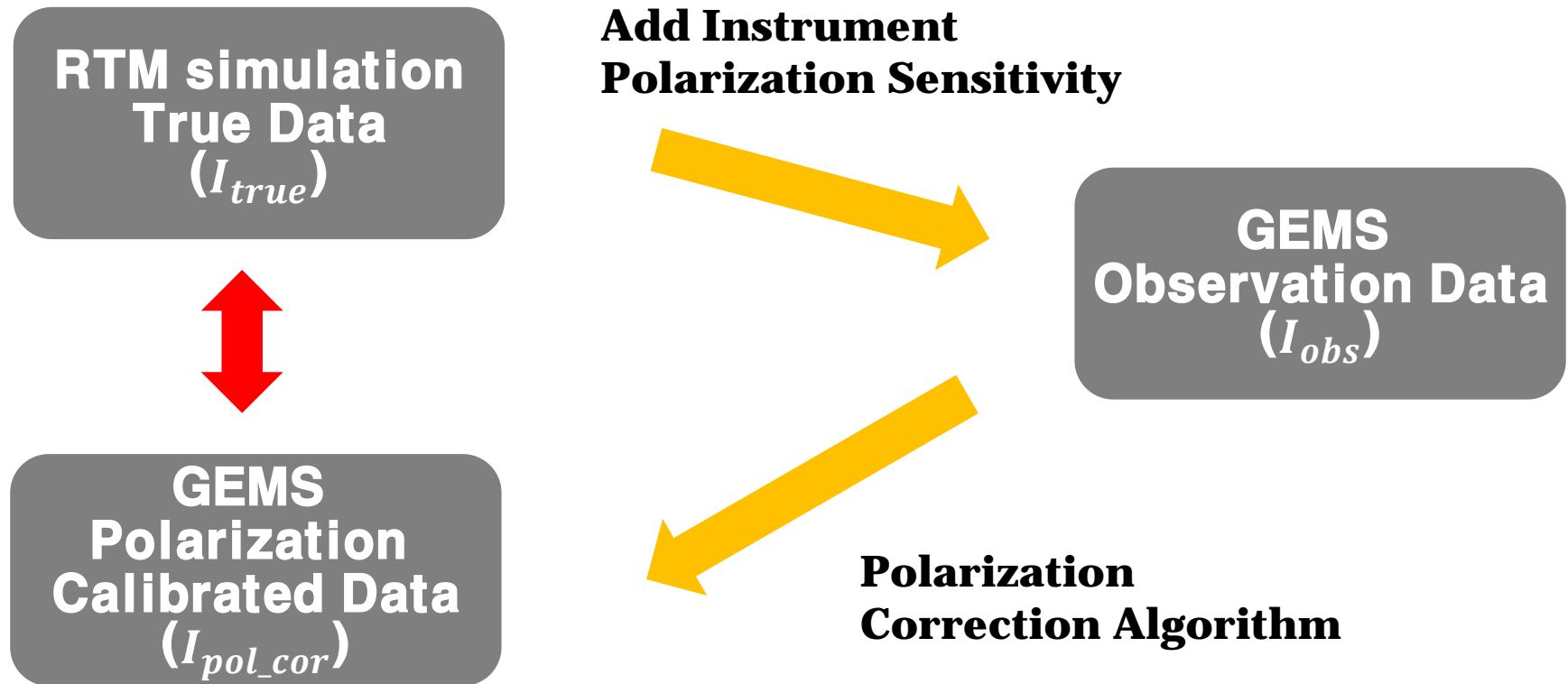


Q/I Bias ( GOME2 - RTM )



- ❖ Simulation of stokes fraction(Q/I) for Rayleigh atmosphere.
- ❖ Observation and RTM simulation are in good agreement.
- ❖ Large differences are observed in the cloud pixels.
- ❖ Polarization correction for clouds might reduce error.

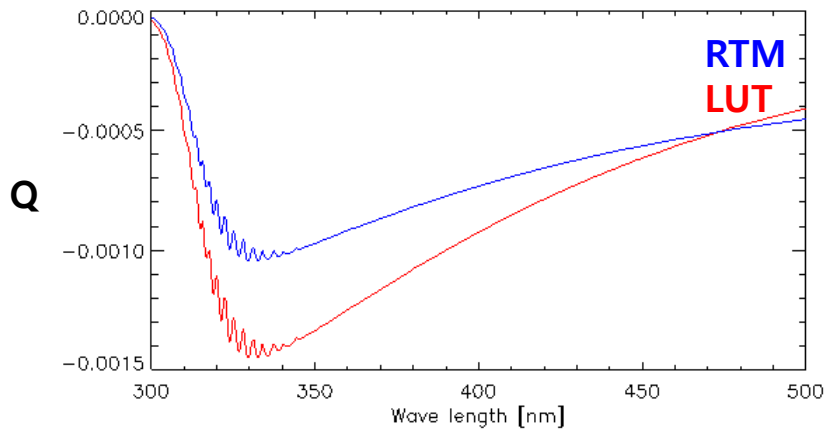
# Test for Synthetic Data



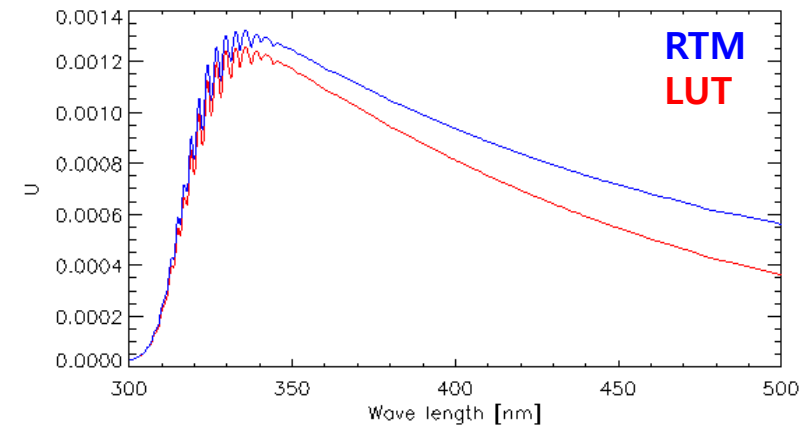
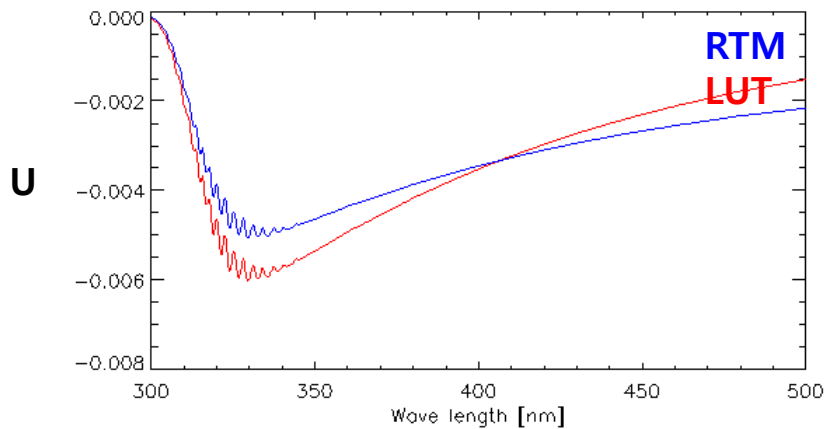
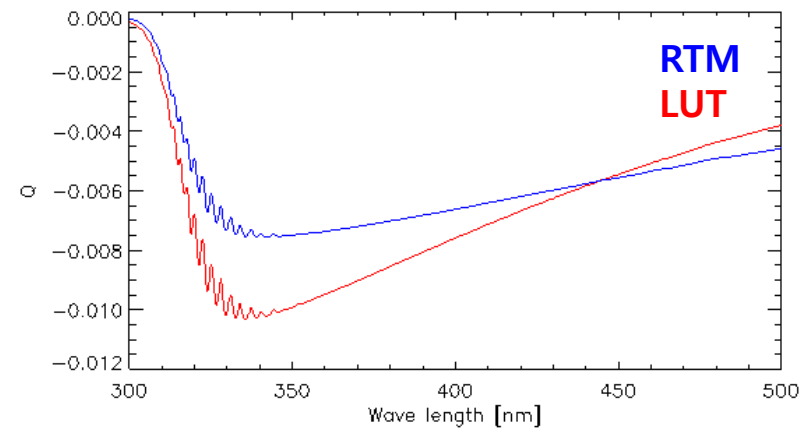
- ❖ Generated GEMS data ( $I_{obs}$ ) from RTM simulation data ( $I_{true}$ ) by adding the instrument polarization sensitivity.
- ❖ The GEMS polarization correction algorithm using LUT was applied to get corrected radiance ( $I_{pol\_cor}$ ).

# Comparison of RTM and LUT

N/S pixel = 500



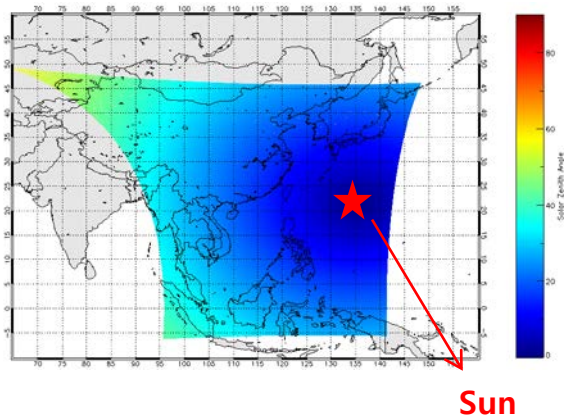
N/S pixel = 2000



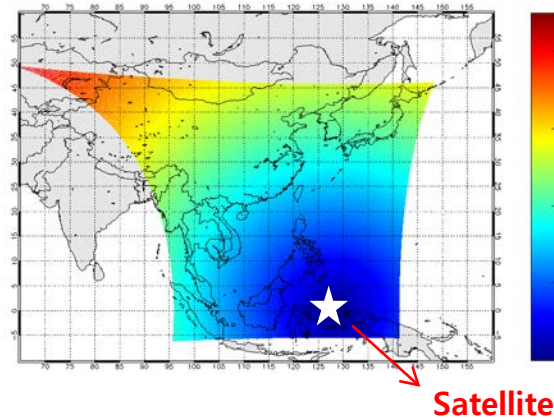


# Synthetic Data Test (20130715 03UTC)

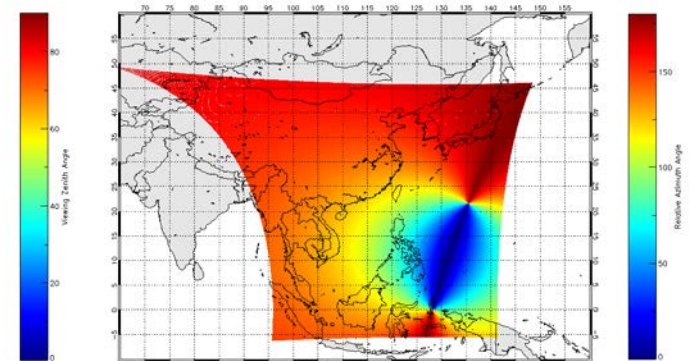
## Solar Zenith Angle



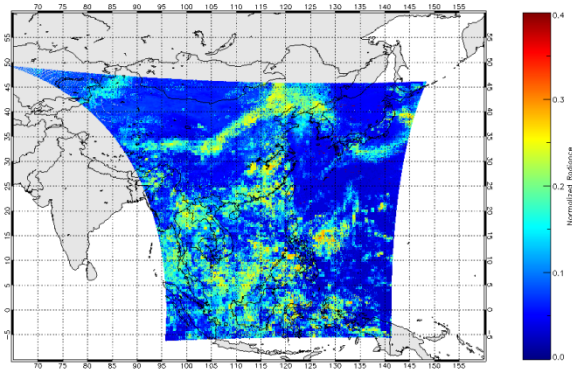
## Viewing Zenith Angle



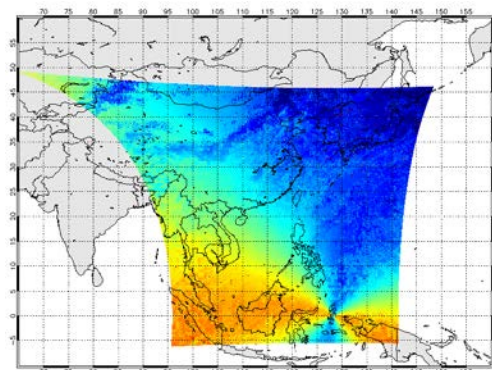
## Relative Azimuth Angle



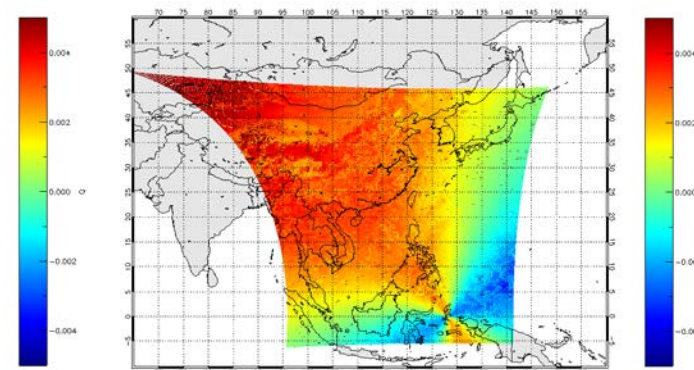
## Normalized Radiance



## Q

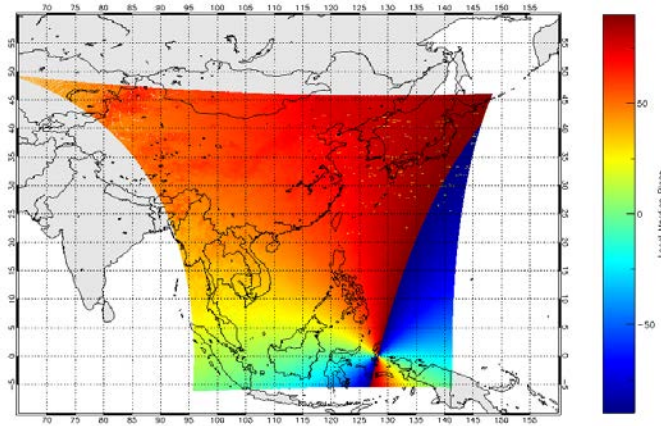


## U

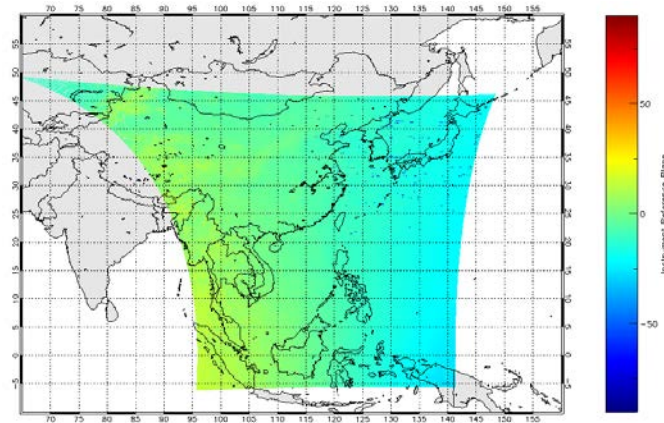


# Synthetic Data Test (20130715 03UTC)

**Polarization Angle ( $\chi_{LMP}$ )**  
Local Meridian Plane

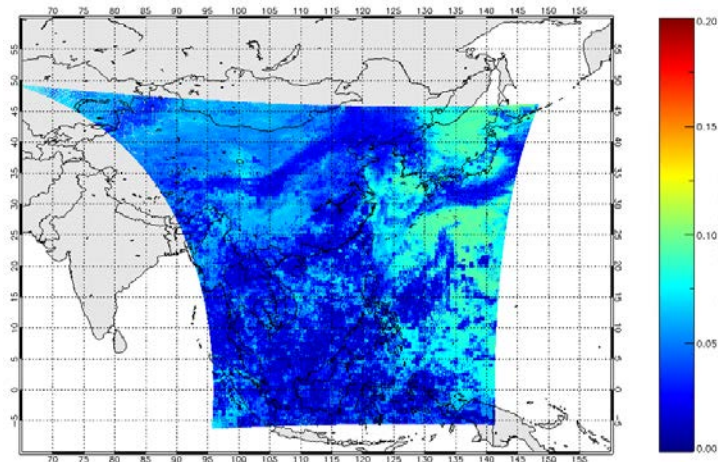


**Polarization Angle ( $\chi_{IRP}$ )**  
Instrument Reference Plane



$$\chi_{LMP} = \frac{1}{2} \arctan\left(\frac{U}{Q}\right)$$

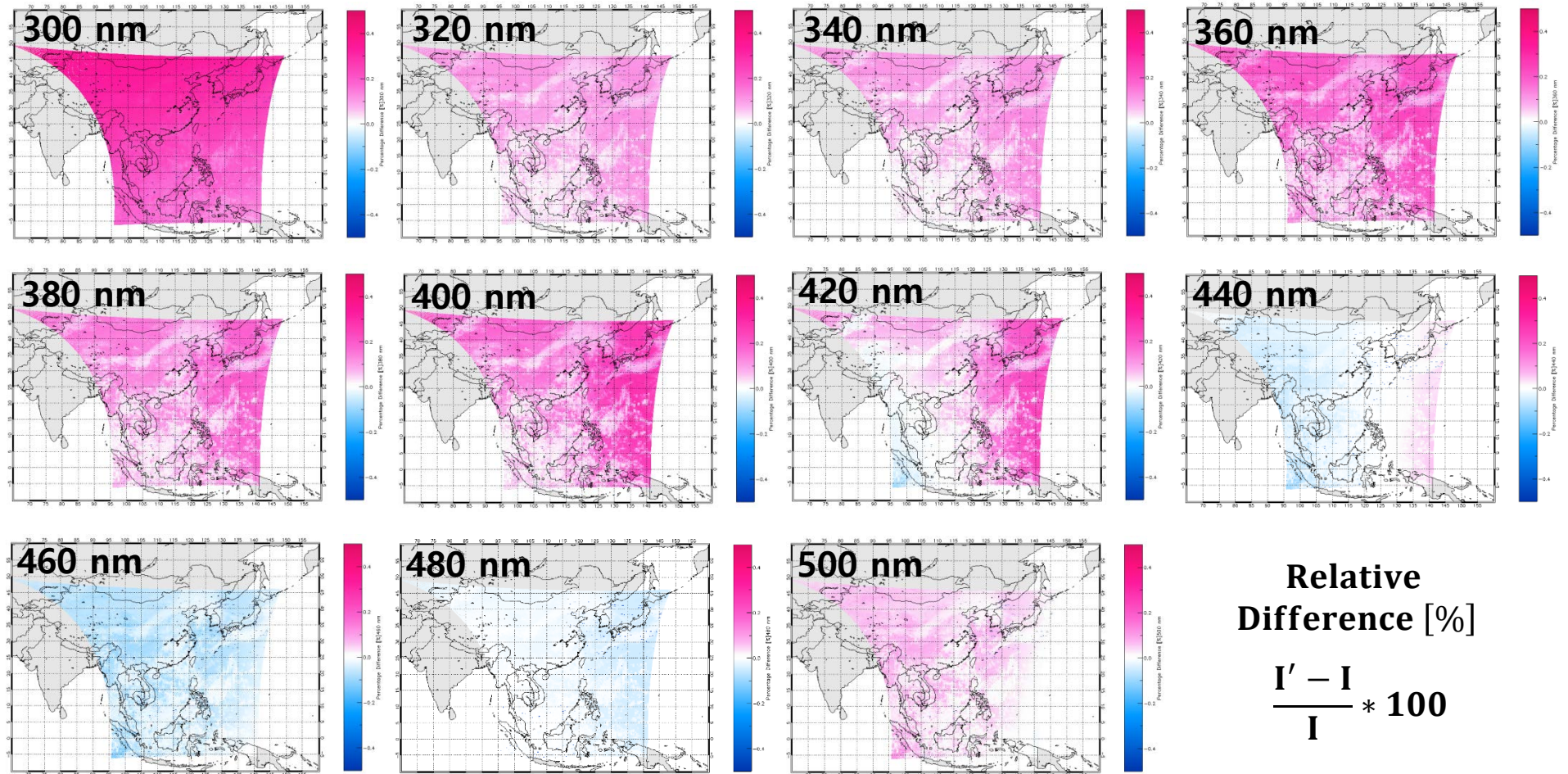
**Degree of Linear Polarization ( $a$ )**



$$\chi_{IRP} = \chi_{LMP} + \Delta\chi$$



# Synthetic Data Test (20130715 03UTC)



Relative  
Difference [%]

$$\frac{I' - I}{I} * 100$$

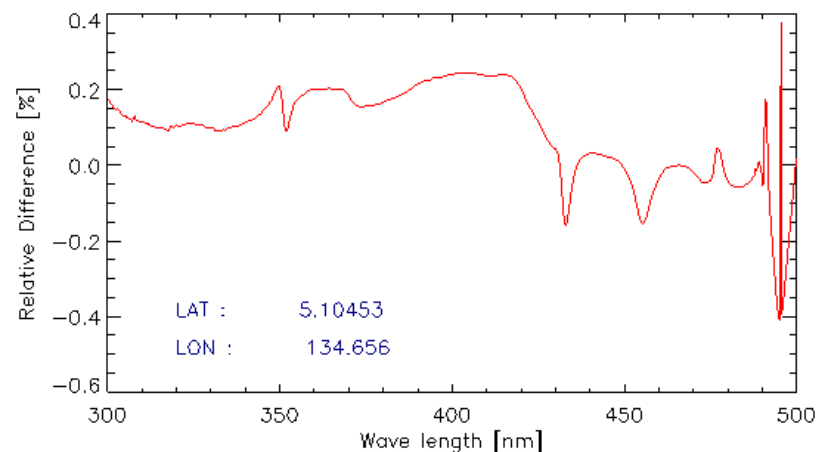
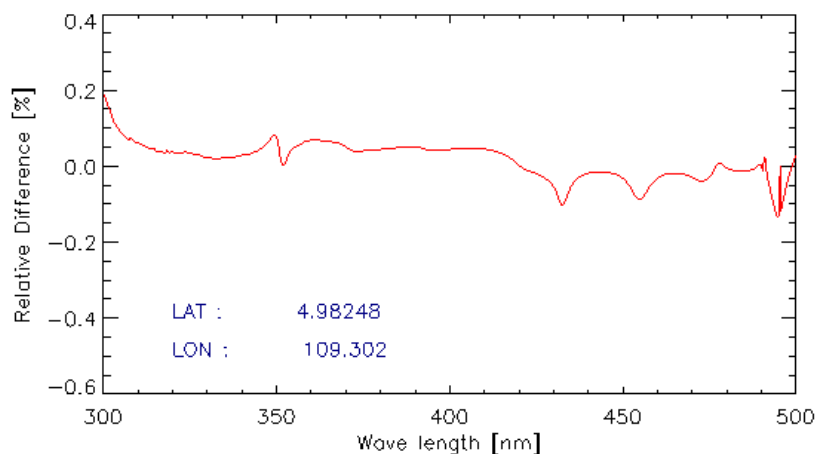
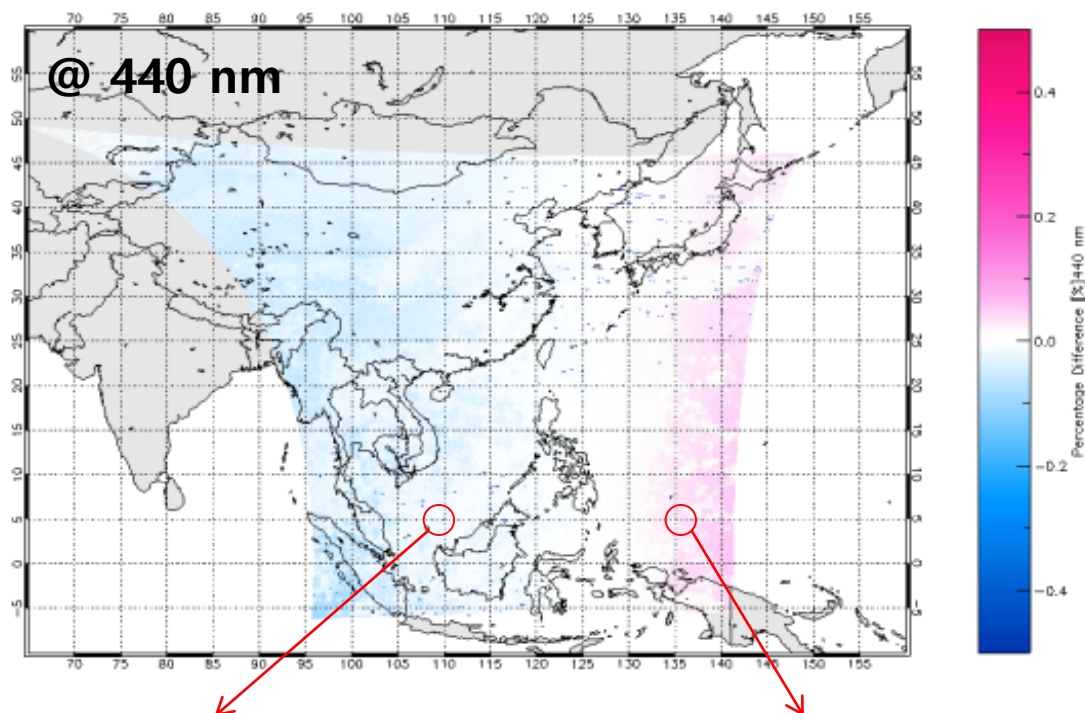
❖ The relative difference between  $I$  and  $I'$  depends on the observation geometry (SZA, VZA, RAA) and wavelength.

# Synthetic Data Test (20130715 03UTC)

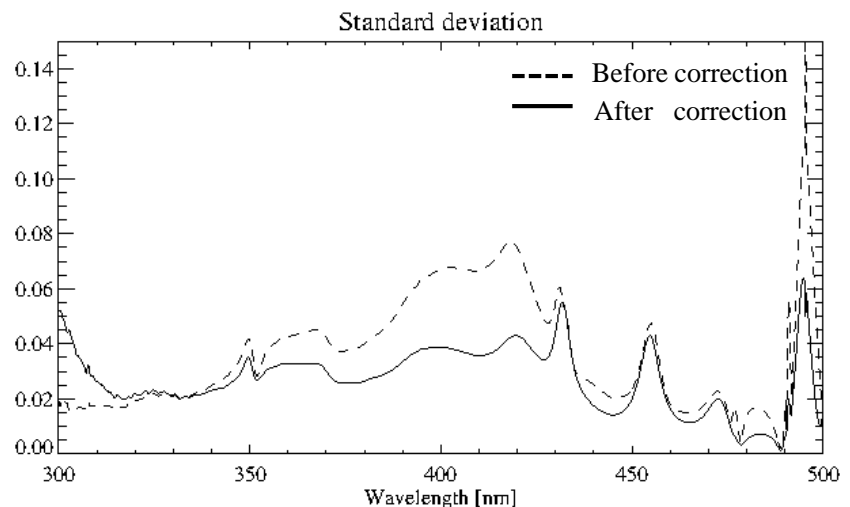
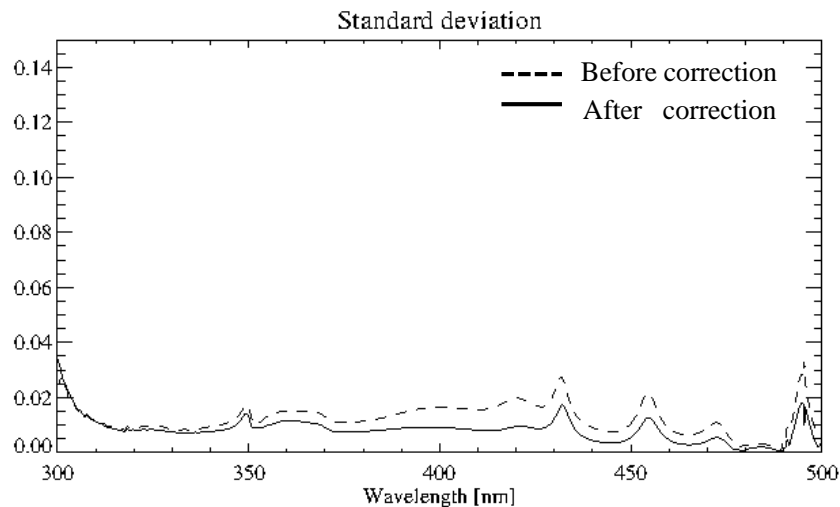
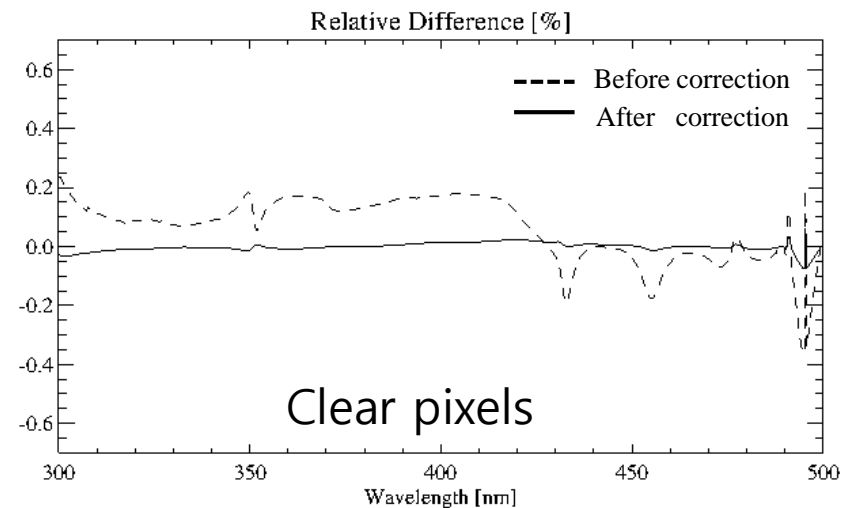
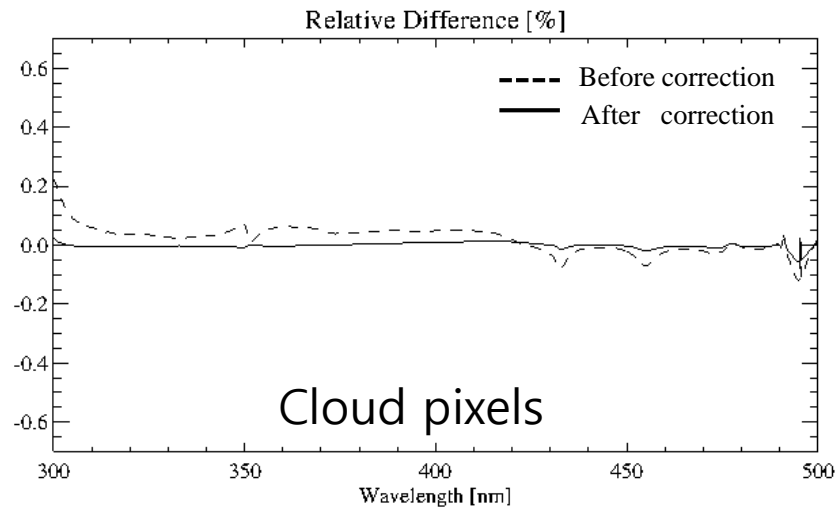
Relative Difference [%]

$$\frac{I' - I}{I} * 100$$

- ❖ The shape of the relative difference depends on the LPS and PA as well as SZA, VZA, ... etc.
- ❖ In these pixels, the relative difference is up to 0.4 %.



# Effects of Polarization Correction



Before :  $(I_{\text{obs}} - I_{\text{true}}) / I_{\text{true}}$

After :  $(I_{\text{pol\_cor}} - I_{\text{true}}) / I_{\text{true}}$

# IOT and Future Plan

- ▣ **Verification of Polarization Correction during IOT**
  - Comparison of GEMS data with RTM simulation for target scenes with known meteorological and chemical field  
(e.g. clear, desert, ocean and opaque convection cloud)
  - Inter-comparison with other satellites  
(e.g. TROPOMI, Sentinel-5 and etc.)
  
- ▣ **Optimization of algorithm ( Accuracy and Speed )**
  - Improve Look-Up Tables
  - Correction for cloud scenes



# Conclusion

- ❑ Polarization characteristics of atmosphere were pre-simulated using RTM for GEMS polarization correction.
- ❑ Polarization error depends on the observation geometry, trace gases, surface information, and etc.
- ❑ Improve accuracy of GEMS L1B data through polarization correction, incorporating instrument polarization characteristics.
- ❑ During the IOT period, we will evaluate and optimize the polarization correction algorithm.



**Thank you ~**

